

Experimental cross sections of the radiative proton-capture reaction $^{112}Cd(p,\gamma)$ ^{113}In inside the Gamow window

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Introduction and Motivation

Importance:

- 113In is generally considered a p nucleus, known to be significantly underproduced in most astrophysical models
- Measurements of reaction rates and cross sections in this mass regime provide stringent tests to the theoretical models
- Measurements of cross sections inside the Gamow window (1.8-4.5 MeV) are expected to provide data for better understanding the *p* process in this mass region

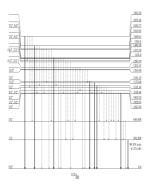
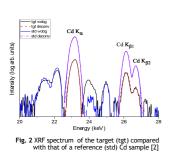


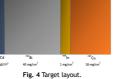
Fig.1: Partial level scheme of the lowlying energy states of ¹¹³In. Solid arrows correspond to transitions feeding the g.s., which were observed in the spectra [1]

The Target

- An isotopic ¹¹²Cd (enriched 97.7%) layer, evaporated on a ^{nat}Bi layer, backed by a ^{nat}In and a thick ^{nat}Cu layer was used. The target thickness was measured via:
- Rutherford Backscattering Technique (RBS)
- > X-Ray Fluorescence (XRF) [2]







Experimental details

- The reaction ¹¹²Cd(p, y)¹¹³In was studied at beam energies of 3.4, 3.5, and 3.6 MeV provided by the 5.5 MV Tandem Accelerator at NCSR "*Demokritos*"
- The subsequent y decays were detected using an array of three HPGe detectors at 55, 90 and 165 degrees, respectively
- The induced neutron-emission channel was additionally studied



Fig. 5 The 5.5 MV Tandem accelerator at NCSR "Demokritos".

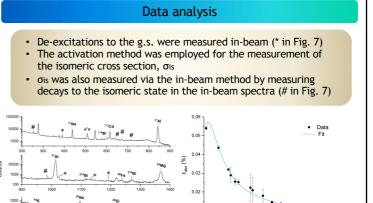
References

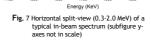
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A. Koning et al., TALYS User Manual, NRG, The Netherlands, 2013



Fig. 6 The array of HPGe detectors.









E [KeV]

Fig. 8 Typical absolute efficiency curve for the detectors employed in this experiment.

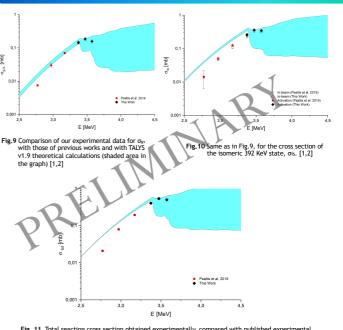


Fig. 11 Total reaction cross section obtained experimentally, compared with published experimental data from previous works of our group [2] and with a full range of TALYS v1.9 theoretical calculations (shaded area in the graph) [1]. Please note that the old and new measurement at 3.4 MeV overlap fully.

Discussion and Future Directions

- The present experimental data confirm recently published data from our group [2] and extend the results to energies above the neutron emission threshold
- Hauser-Feshbach calculations with TALYS agree well with the results for the g.s., but overestimate the results for the isomeric state
- Additional measurements to obtain the astrophysical S-factor and the nuclear reaction rate in this energy region are underway
- \cdot Future work involves detailed calculations with the most recent TALYS v1.95 code to improve our knowledge of the OMP, NLD and γSF in this region

Acknowledgments

AK and AZ acknowledge support by the Hellenic Foundation for Research and Innovation (HFRI) and the General Secretariat for Research and Technology (GSRT). The NuSTRAP group at U. Athens acknowledge support by ENSAR2.

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